EXPERT AE SIGNAL ARRIVAL DETECTION

Milan Chlada

Institute of Thermomechanics AS CR, Czech Republic

ABSTRACT

Accurate acoustic emission (AE) source location is the primary goal of the defect analysis following the AE signal detection. The source localization is mostly based on arrival time differences of signals recorded by several transducers. Considerable signal distortion happens during the wave propagation through the solid. Inaccurate determination of signal onset and arrival time differences respectively, are the greatest sources of localization errors. Especially, in a case of higher requirements on accuracy and robustness, the results of currently used localization methods appear to be insufficient. In the paper, recently improved version of the new signal - shape based algorithm, modeling an expert system of the elastic wave arrival detection, is introduced. In many applications, this method, based on signal energy and local gravity center evolution, has been proved as rugged enough, fast and easily applicable.

KEYWORDS

Acoustic emission, signal arrival detection.

1. Introduction

In present, there exist various approaches to the AE signal arrival detection [1] such as e.g. the threshold crossing method, the first signal onset detection, time differences detection by means of cross-correlation function, and the methods outgoing from the wavelet transformation, or detecting local statistical parameters changes of the signal in terms of time series [2]. However, with respect to accuracy, automation and robustness, these methods often fail to satisfy more exacting requirements resulting from the large and variant AE signal sets processing needs. A number of algorithms also demand to resize input parameters even more times within one measurement data analysis (e.g. due to the different channels characteristics of measuring device).

For the fast and rugged signal processing needs, the new signal edge detection algorithm was proposed. The method should reach relevant results for AE signals of various complex shapes, different noise levels, partially broken, highly distorted or incomplete (see fig.1). The algorithm goal is to simulate an expert ("visual") elastic wave arrival time detection based on local gravity center or signal energy evolution.

Fig.1: typical AE signal shapes

2. Algorithm description

2.1. Measuring window length computation

The unique algorithm input is sampled signal $s = \{s(j) | j=1, \dots, N\}$. First of all, the window length adjusting parameter k is computed. To suppress the local gravity center (2.5) vector oscillation, maximum window width ($=2k+1$) should slightly exceed the "main" data record frequency wave length, i.e. double maximum length between the neighbouring signal mean value crossings. The measure of "exceeding" is represented by parameter c_k , which should not be, indeed, too extensive. Unlikely, the method accuracy would decrease. Parameter k is computed as follows:

$$k = c_k \cdot \max(\text{diff}(z)) \quad (2.1)$$

Where the auxiliary vector $z = \{z(m) | m=1, \dots, N_m\}$ is made up by N_m ascendantly sorted signal sample indexes, after which zero crossing occurs (when the signal is already centered):

$$z(m) = j_m \text{ pro } \forall m: 1 \leq m \leq N_m \wedge |d(j_m)| = 2 \wedge 1 \leq j_m < j_{m+1} < N, \text{ where } d = \text{diff}(\text{sign}(s)) \quad (2.2)$$

The difference operation is defined as distinction between original vector samples:

$$\text{diff}(x) = \{ x(j+1) - x(j) \mid j=1, \dots, N-1 \} \quad (2.3)$$

2.2. Auxiliary local gravity center vector

Subsequently, for all signal samples, the auxiliary vector g is computed. It's i -th element is the gravity center y-coordinate $GC_Y(|s_i|)$ of existing absolute signal part s_i around the relevant sample i :

$$g(i) = GC_Y(|s_i|) = \frac{\sum_{j=1}^{k_i} (k_i+1-j)|s_i(j)|}{\frac{1}{2}(k_i^2 + k_i)} \quad (2.4)$$

Where above mentioned existing signal part is considered as:

$$s_i = \{ s(j) \mid j = \max\{i-k, 1\}, \dots, \min\{i+k, N\} \} \quad (2.5)$$

2.3. Automatic noise separation and approximate signal edge appointment

The key problem of designed automatic algorithm is an approximate signal noise part separation. Considering signal samples as the unique input data, it is needed to propose special functions providing distinct amplitude increase detection and, afterwards, noise pretrigger removal. Due to dispersion and various reflections, the steepest signal rise many times doesn't precede maximum amplitude. Further, one data record can represent more overlapping signals and, therefore, maximum amplitude doesn't need to belong to first event, that is preferred (see fig.1 - type2).

Let's assume that the initial record samples are realization of a random process with constant mean value and variance. Unlike the threshold detection, the new algorithm should not respond to solitary amplitude deviations as e.g. electrical interferences or AD converter failures (see fig.1 - type2). In practice, we often meet situations when the signal beginning is captured just very closely by the first record samples or actually anywise. Automatic algorithm should solve even those situations and by means of mathematically modeled expert knowledge detect signal noise part absence or the event record incompleteness. In terms of above mentioned requirements, following functions q_1 and q_2 were established:

$$q_1(i) = \begin{cases} 0 & i=1 \\ \frac{(i-1)(g(i)+g(1))}{\sum_{j=2}^i (g(j-1)+g(j))} & \text{for } i=2, \dots, N \end{cases} \quad (2.6)$$

Function q_1 compares the surface below the join of the first and i -th local gravity centers vector sample with the sum of surfaces below each sample joins in analogical interval (see fig.2). This ratio detects the first marked gravity centers auxiliary vector rise and, therefore, also significant rise of the signal itself.

Vice - versa, it doesn't reflect unique signal deviations. Relative algorithm stability is provided in this way,. The peak sample index m_1 of vector q_1 determines the first distinct rise point of the signal .

Fig.2: vector q_1 computing

Fig.3: vector q_2 computing

Before function q_2 implementation, it is necessary specify highest local gravity centers value index m_2 in the range till distinct rise point m_1 . Generally, these numbers m_1 and m_2 can differ.

$$m_1 = \underset{i}{\operatorname{argmax}} \{q_1(i)\} , \quad m_2 = \underset{i \leq m_1}{\operatorname{argmax}} \{g(i)\} \quad (2.7)$$

Conversely, function q_2 compares the sum of surfaces below each local gravity center vector samples joins between points i and m_2 with the surface under relevant interval outer samples join (see fig. 3). Such ratio rapidly increases at the point of signal amplitude elevation over noise background:

$$q_2(i) = \frac{\sum_{j=i+1}^{m_2} (g(j-1) + g(j))}{(m_2 - i)(g(m_2) + g(i))} \quad \text{for } i=1, \dots, m_2-1 \quad (2.8)$$

By reason of the noise characteristics estimate it is necessary to reflect initial emission event record samples, which do not interfere to effective signal, naturally. This noisy part can be specified by the point z_s , where the function q_2 crosses certain limit p_z . Actually, algorithm debugging showed that in incomplete beginning signal record case function q_2 starts with higher value than limit p_z . According to numerical experiments results, optimum p_z value lies in range 0.2-0.3:

$$z_s = \min \{ i \mid q_2(i) > p_z \} \quad (2.9)$$

In case of record noise part, also local gravity center vector values are almost constant, analogous to statistical signal characteristics. Hence, the signal sample location, where auxiliary vector g lastly exceeds the noise characterizing level, can be considered as an approximate AE signal arrival. Such as threshold level p_s , the maximum local gravity center vector value into sample z_s is taken:

$$p_s = \max \{ g(i) \mid i \leq z_s \} \quad (2.10)$$

Approximate arrival z_a is finally obtained by following thresholding:

$$z_a = \max \{ i \mid g(i) < p_s \} \quad (2.11)$$

Fig.4: signal noise part determination

2.4. Measuring window length factor

Measuring window length is important accuracy affecting factor in relation to signal record noise level. In hypothetical ideal case, when the noise background is not present, immediate signal amplitude influences computed local gravity centers vector by k samples earlier, then the currently evaluated element is. Conversely, if the signal beginning is hidden by noise, gravity centers vector change will approve more gradually. After many numerical experiments results, the measuring window length factor f_w was established to compensate this effect and to correct the estimated signal beginning position with the linear reference to the ratio of noise level and local gravity centers vector maximum. The factor's signal - to - noise ratio sensitivity is adjusted by constant c_w :

$$f_w = c_w \cdot \frac{p_s}{\max \{ g(i) \}} \quad (2.12)$$

2.5. Signal arrival estimation

Signal record noise part is relatively exactly specified by approximate arrival z_a . In this way provided noisy segment and it's characteristics are exploited in final algorithm improvement steps. As the auxiliary threshold level p_1 , the sum of relevant gravity centers vector part mean value and standard deviation is considered:

$$p_1 = \text{mean} \{ g(i) \mid i \leq z_a \} + \text{std} \{ g(i) \mid i \leq z_a \} \quad (2.13)$$

The last vector $g(i)$ sample, that is less than value p_1 , is rated as first signal onset estimate (see fig.5) and subsequently corrected by measuring window length factor:

$$z_1 = \max \{ i \mid g(i) < p_1 \} + k \cdot f_w \quad (2.14)$$

2.6. Correction to zero crossing

Each signal arrival estimate can be put more precisely by correction to zero crossing. It concern finding the last place before the actual arrival estimate location z_{act} , where the signal virtually comes through the mean value. The sample on the right from this place is taken as the new signal beginning version z_n :

$$z_n = \max \{ i \mid i \leq z_{act} \wedge |d(i)| \neq 0 \}, \text{ where } d = \text{diff}(\text{sign}(s)) \quad (2.14)$$

All introduced signal arrival versions are always revised to “zero crossing”.

Fig.5: vector $g(i)$ thresholding

Fig.6: signal arrival linear extrapolation

2.7. Linear extrapolation of signal arrival

Particularly in cases of high signal noise levels, exact arrival estimate is more complicated. Using such oscilloscope data record, the signal onset recognition is not unique even by naked eye. This problem is illustratively shown on fig.6 with typical record and the same one with high level white noise added. It is evident that original signal arrival is entirely covered by noise and the detection method lost it's exactness (see fig.6 - arrival version z_1). Hence, the signal arrival linear extrapolation was proposed. This elementary solution helps also in case of incomplete data record. Due to the general signal's linear initial amplitude rise, the first onset detection improvement can be done by the linear extrapolation marked on fig.6. The line comes through the points with coordinates $[m_2, g(m_2)]$ and $[z_1, g(z_1)]$. X-coordinate intersection z_2 estimates the signal arrival more exactly.

Evidently, when the signal noise level is low, such arrival location extrapolation z_2 basically keeps the position z_1 unchanged. If the version z_1 is located near the record start, resulting value z_2 can be even negative. Thereby, certain not registered signal start is automatically restored. The benefits of this extrapolation were evidently proved by many numerical experiments.

3. Other algorithm versions

Regarding very similar features, certain method alternative can be represented by analogical local RMS values finding. All algorithm steps may be unchanged, except

the auxiliary vector $g(i)$ (originally see 2.4) should be computed as the relevant signal part s_i of RMS parameter:

$$g_{RMS}(i) = RMS(s_i) = \sqrt{\frac{1}{k_i} \sum_{j=1}^{k_i} (s_i(j))^2} \quad (2.15)$$

Next algorithm steps, i.e. computing of functions q_1 and q_2 , thresholding, and arrival estimate corrections, remain similar. However, unlike local gravity centers version, parameter RMS approved itself somewhat sensitive to immediate signal amplitude and this method modification looses robustness against various interferences in data record. Obversely, during numerical testing, RMS parameter proved better in higher noise level cases.

Within noise development observation by auxiliary vectors g a g_{RMS} it is possible to establish an alternative threshold values p_{alt} , e.g. the noise part maximum or mean of both auxiliary vectors g and g_{RMS} , including partial GC and RMS of relevant vectors g and g_{RMS} respectively:

$$p_{alt} = \max / \text{mean} / GC / RMS \{ g_{(RMS)}(i) \mid i \leq z_a \} \quad (2.16)$$

4. Parameters adjustment

Algorithm parameter values optimization was made using the set of 150 randomly selected AE signals measured during complex aircraft structure part pen-test excitation. For the method precision examination, not only the original high signal - to - noise ratio records were used, but also similar ones with the white noise added. The estimative optimal parameter values were obtained by sequential approximation during numerical experiments without analytic computation. Each arrival estimation method has different proper parameter settings, naturally. Nevertheless, for given testing data set, the optimal values are in the neighbourhood of next ones:

$$c_k \simeq 1.3, \quad p_z \simeq 0.25, \quad c_w \simeq 8 \quad (2.17)$$

With above mentioned adjustment, the new algorithm achieves significantly better results (the best mean absolute error was 18 samples) than edge detection method (mean absolute error was 199 samples). Considering appropriate sampling frequency (200ns) and ultrasound pulse propagation velocity (6.18mm/us), we obtain resulting AE event localization improvement from 246 to 103mm.

5. Algorithm acceleration

Whole signal length auxiliary vectors computing can be quite time - consuming and basically useless. For signal development casual information there is no need to calculate all samples of auxiliary vectors g a g_{RMS} . However, it is fully sufficient to shift measuring window within half of it's length, thus, by k samples. Such as resolution is enough for distinct rise m_2 estimative determination. Indeed, after that it is necessary to recalculate auxiliary vectors in full resolution, nevertheless, just from the point where vector q_2 exceeds the value about 0.21 (signal noise part mostly gradually ends here) to sample m_2 . Considerable numerical operations amount is saved this way, so the algorithm is much faster and effectually applicable.

6. Conclusions

In contribution, the new AE signal arrival detection algorithm is described. Main features of the method can be summarized into following points:

- Emission event signal arrival is determined in terms of auxiliary local gravity centers and energy (RMS) vectors analysis.
- Suitably sampled signal is the unique algorithm input. An automatic measuring window width adjustment and signal record noise part estimation is implemented. Newly designed internal functions reliably separate the noise through whole it's approximate length, not only priorly set partition.
- New method is featured by robustness and reliability. It is debugged using real experiment extensive data set. The electrical interferences and AD converter failures sensitivity is highly suppressed.
- Algorithm parameters were numerically optimized. The accuracy achieved is markedly better then that one of up to now applied methods.
- In incomplete signal record detection and higher noise level cases, signal arrival linear extrapolation is performed, whereby the noisy or missing signal initiation is reconstructed.
- Optimized algorithm implementation in Matlab provides sufficiently fast commonly processed AE signal records analysis.

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